

Basic Physics Problems

[A] Basic Physics : Classical Mechanics

- (1) Consider the motion of planet with mass of m in the gravitational potential together with the additional potential of $\frac{\alpha}{r^2}$. Note that the additional potential is not integrable. The total gravitational potential becomes

$$U(r) = -\frac{GMm}{r} + \frac{\alpha}{r^2} \quad (1)$$

where M denotes the mass of the sun.

- (i) Find an exact solution of the Newton equation.
 - (ii) Prove that the orbit should have a discontinuity.
- (2) The Newton equation with non-integrable potential must be solved in a perturbation theory.
- (i) Find a perturbative solution of the planet orbit.
 - (ii) Explain the behavior of orbit perihelion shifts.

[B] Basic Physics : Electromagnetisms

- (1) Describe some physical properties of Poynting vector.
- (2) Is the Poynting vector related to the electromagnetic wave or not ?
 - (i) If yes, then explain its reason.
 - (ii) If no, then explain the physical picture why the Poynting vector is not related to photon.
- (3) Explain any physical reasons why the vector potential must be quantized.
- (4) Dirac field must be quantized with the anti-commutation rule. Explain the physical reason of this quantization method.
- (5) The equation of motion of free photon can be written as

$$\partial_\mu F^{\mu\nu} = 0, \quad \text{with} \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad (2)$$

where $F^{\mu\nu}$ denotes the electromagnetic strength. In this case, the number of freedoms of vector potential A^μ is four while photon has only two. Explain why the two degrees of freedom are lost.

[C] About the right-hand side of Einstein equation.

- (1) Suppose the state vector is denoted by $\Psi(x)$. In this case, write explicitly the energy-momentum tensor $T^{\mu\nu}$.
- (2) In classical mechanics, the energy-momentum tensor $T^{\mu\nu}$ cannot be defined. Explain this reason.
- (3) In Einstein equation, the energy-momentum tensor $T^{\mu\nu}$ is defined. Answer as to why it can be defined.
- (4) Questions for star formation and star distribution function.
 - (i) Explain any interactions that determine the star formation and star distribution function.
 - (ii) Describe the properties of these interactions.
 - (iii) For the star formation, which interaction may play what role?
 - (iv) For the star distribution, which interaction may play what role?
- (5) Black Hole is not defined as a star. Explain why people believed that the Black Hole is a star.